

# At-wavelength Alignment and Testing of the 0.3-NA MET Optic

Reaching diffraction-limited imaging, and beyond

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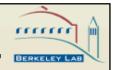
#### LLNL

John Taylor, Don Phillion, Layton Hale, Mike Johnson, Henry Chapman, Nhan Nguyen, Gary Sommargren International SEMATECH Project Manager

Kim Dean



# **Project Goals**



Create and operate an EUV resist-testing facility with imaging down to ~15 nm, and several unique capabilities. (see Naulleau, et al.)

For Optimal EUV imaging, wavefront tolerances are ~0.1 nm

# Ultra-high accuracy *EUV interferometry*

- Many opportunities for learning
- Extensions of known techniques
- Opportunity for cross-comparison



The MET (Set-2)

(shown here with surrogate optics)

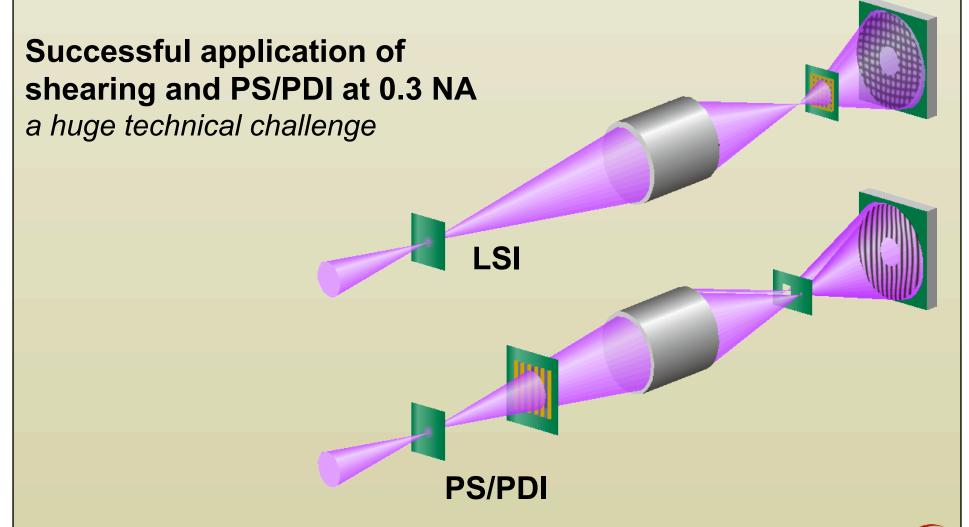
Made by <u>Zeiss</u>.
Assembled and pre-aligned by
<u>Lawrence Livermore</u>.



#### At-wavelength MET-testing overview



EUV interferometry, alignment and characterization



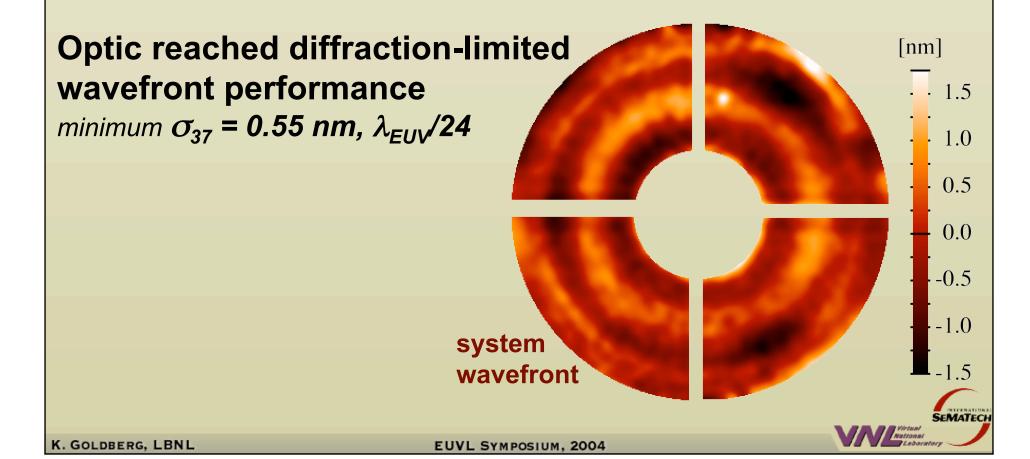


# **At-wavelength MET-testing overview**



EUV interferometry, alignment and characterization

Successful application of shearing and PS/PDI at 0.3 NA



# At-wavelength MET-testing overview



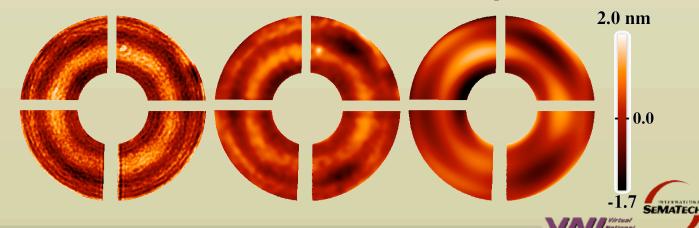
EUV interferometry, alignment and characterization

Successful application of shearing and PS/PDI at 0.3 NA

Optic reached diffraction-limited wavefront performance

#### Visible PSDI • EUV PS/PDI • EUV LSI intercomparison

complicated by alignment issues



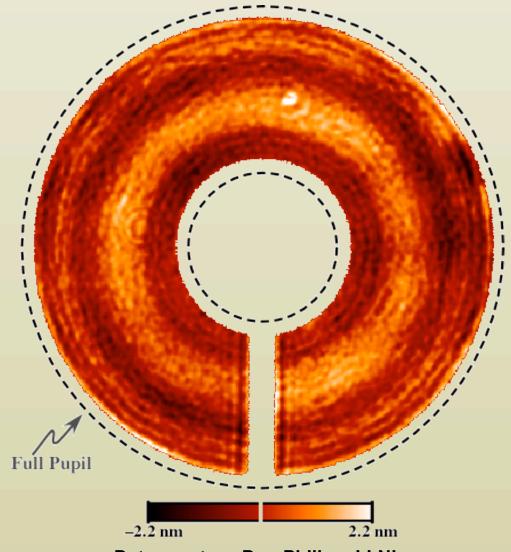
# Three high-accuracy interferometers

$\sim$	Α
umm	[111]
BERKELEY LA	

lensless PSDI	LSI (shearing)	PS/PDI
LLNL Lawrence Livermore	LBNL Lawrence Berkeley	LBNL Lawrence Berkeley
visible-light $\lambda = 532.2 \text{ nm}$	EUV 13.5 nm	EUV 13.5 nm
accuracy target ~λ <sub>vis</sub> /5322	~λ <sub>EUV</sub> /135	< λ <sub>EUV</sub> /135
-Essential for single-element testingConvenient for system alignmentOperates at air.	-Fast, easy to performHigh accuracy requires careful calibration & analysisUsed for field measurement and alignment.	<ul> <li>-The high-accuracy standard.</li> <li>-Working with sub-30-nm pinholes for 0.3 NA testing is a challenge.</li> <li>-Used for accuracy validation and higher spatial-f response.</li> <li>-Covers the full pupil</li> </ul>

#### Final visible-light measurement of the MET





astigmatism, coma, and spherical aberration were "zeroed" by alignment

10 < r < 26 mm

0.56 nm 37-Zernike fit

0.15 nm astigmatism

0.12 nm trifoil

0.10 nm coma

0.05 nm spherical ab.

0.49 nm h.-o. spherical

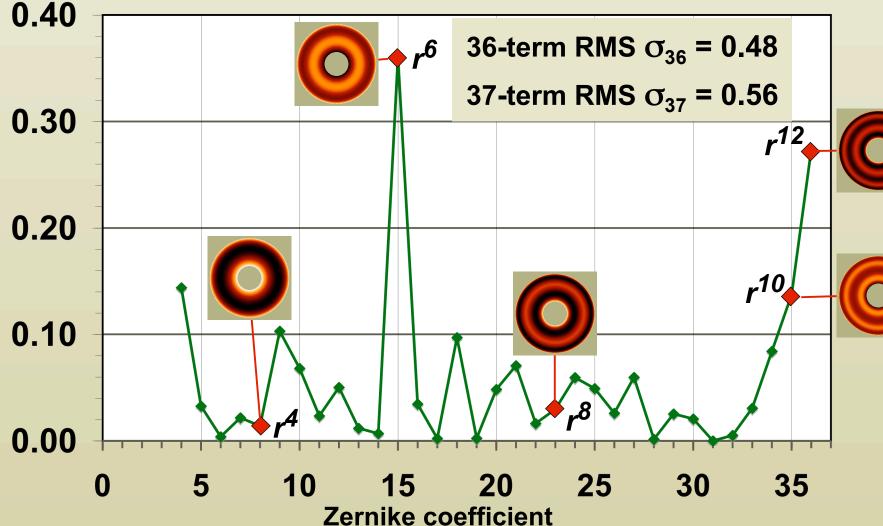
**Data courtesy Don Philion, LLNL** 









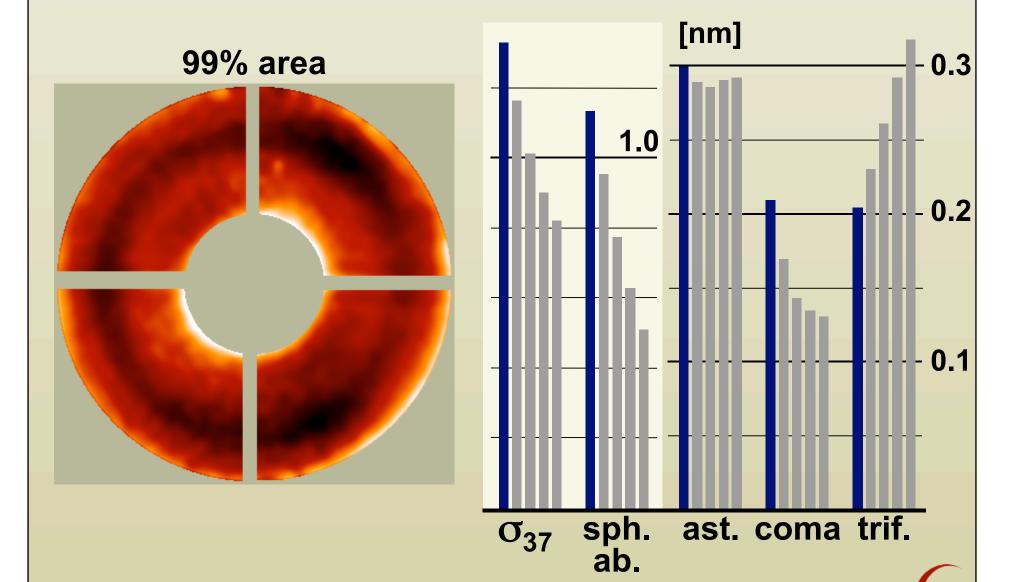


MET Set-2 visible-light data, LLNL/LBNL



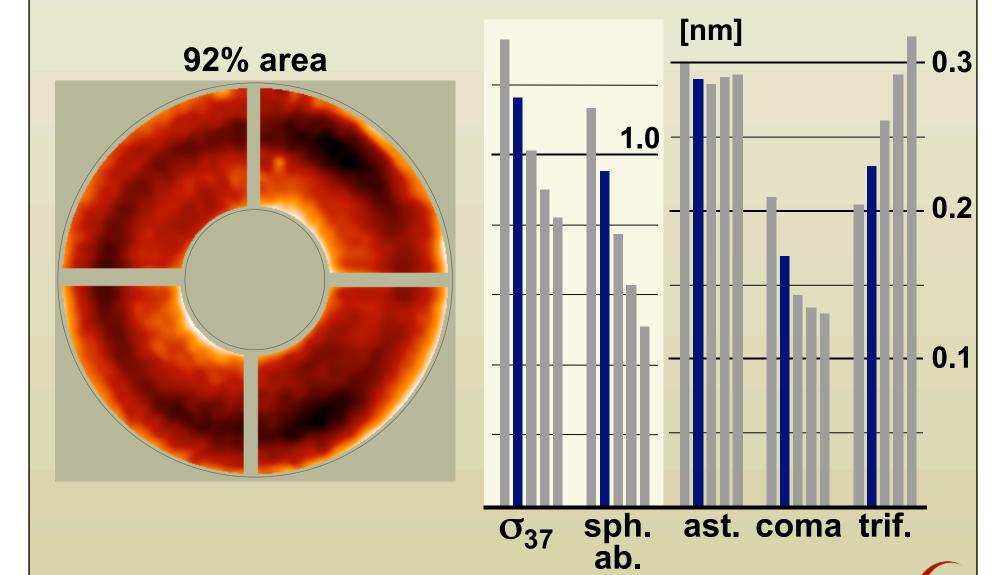






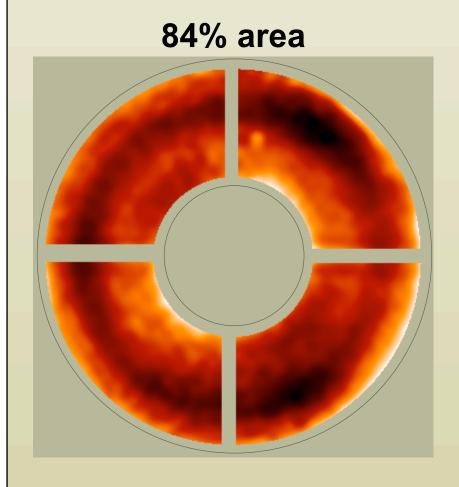


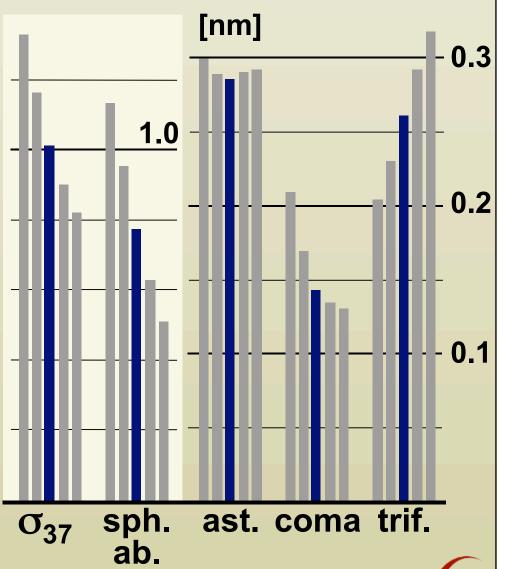






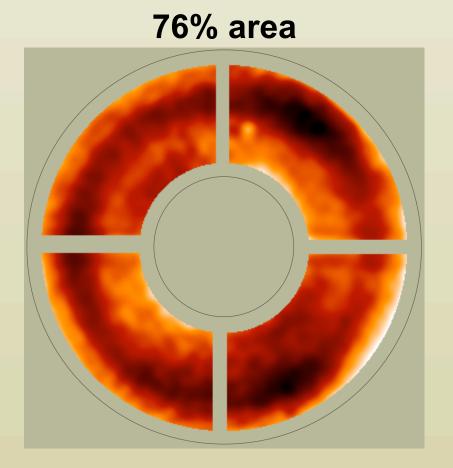


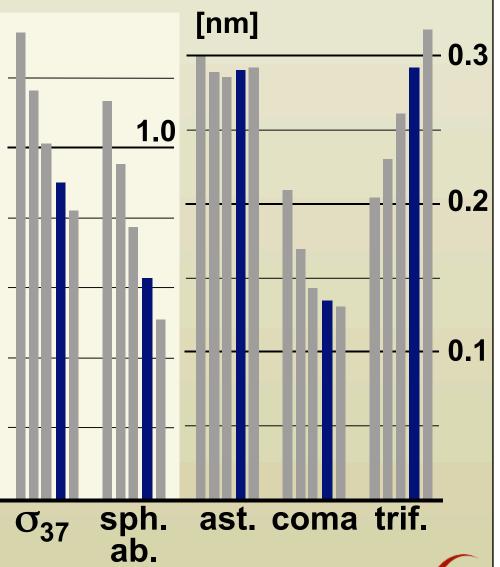






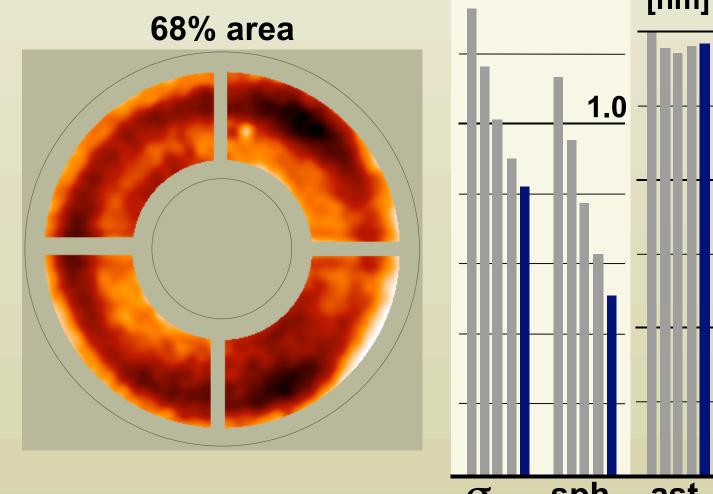


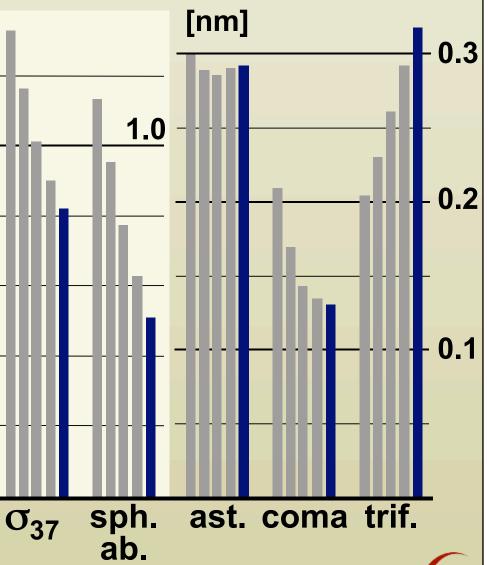




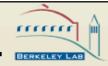


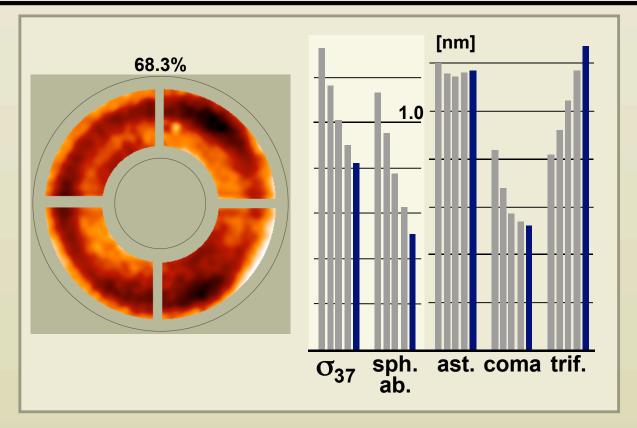






# The importance of measuring the whole pupil





- We cannot predict the aberrations outside of the measurement domain
- Values depend strongly on the pupil area.

Modeling based on only part of the pupil gives you only part of the answer



LLNL



LBNL









one month

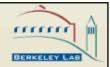


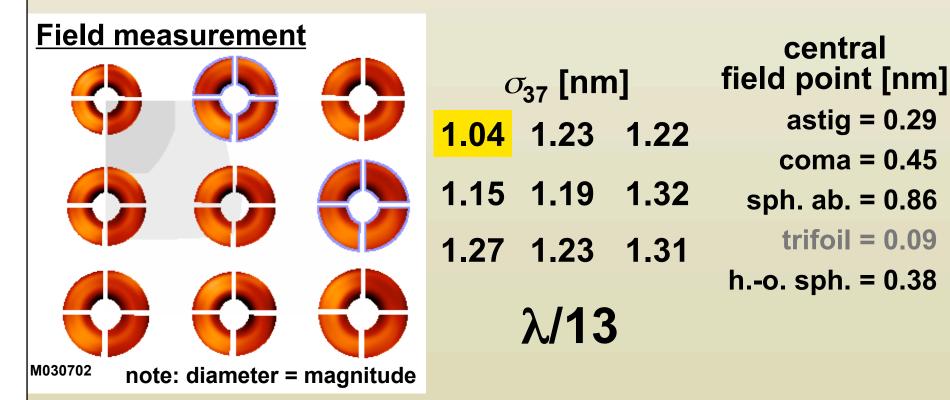




Alignment

#### Initial shearing measurement at 20°C

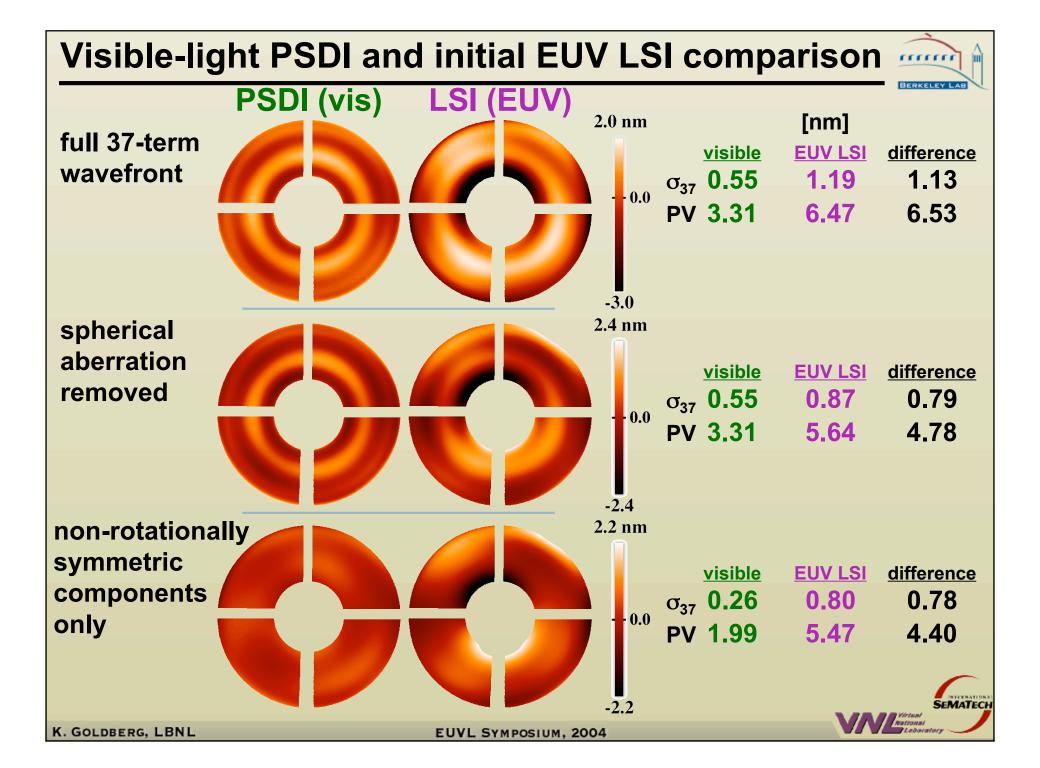




In our first EUV measurements at 20°C, a large, *unexpected* primary spherical aberration was dominant.

Higher-order spherical aberration was also present.





#### First EUV alignment



# Field measurement

O	<sub>37</sub> [nn	1]	central field point [nm]
0.84	0.80	0.66	astig = 0.05
	0.00	0.04	coma = 0.08
0.80	0.66	0.84	sph. ab. = $0.02$
0.94	0.66	0.83	trifoil = 0.22
		_	ho. sph. = 0.34
$\lambda/20.5$			

Astigmatism, coma, and spherical aberration are sensitive to alignment and can be removed.

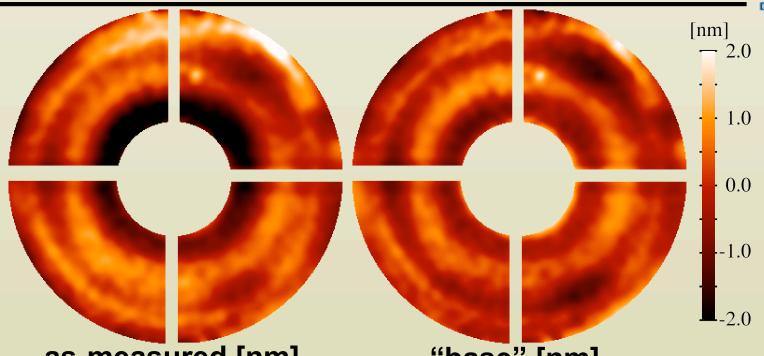
Adjustments are made to M1's 6-arm mount: 0.03-2.50 µm step sizes.



M030806

#### PS/PDI measurements 2 days after LSI alignment





as-measured [nm]

 $\sigma_{37} = 0.68$ 

PV = 3.54

astig. = 0.18

coma = 0.28

sph. ab. = 0.40

**trifoil = 0.10** 

h.-o. s. = 0.30

"base" [nm]

 $\sigma_{37} = 0.45$ 

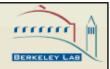
PV = 2.92

With astigmatism, coma, and spherical aberration removed.

The system alignment had changed noticeably in 2 days.



#### System stability



The stability of every optical system is unique.

#### THEORY:

- We believe small alignment actuations contribute to the instability.
- Vent/pump cycles may release stress.
- There is not enough data to draw firm conclusions.

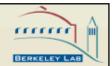
#### **REMINDER:**

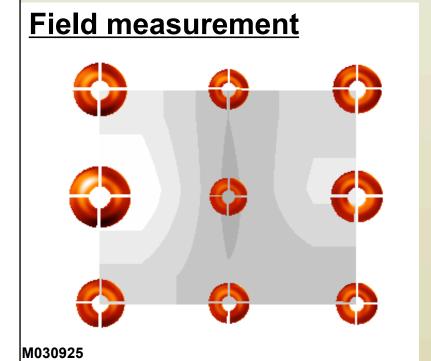
These effects are small, not large.
 Wavefront changes were a few tenths of a nm.

In-Situ Monitoring will be important



# Second (and best) EUV alignment





central field point [nm]  $0.79 \ 0.59 \ 0.71$  astig = 0.04 coma = 0.06  $0.90 \ 0.55 \ 0.76$  sph. ab. = 0.04 trifoil = 0.14 h.-o. sph. = 0.37  $\lambda/24.5$ 

Following the installation of some imaging hardware, the optic was re-measured and re-aligned, achieving its best overall alignment.



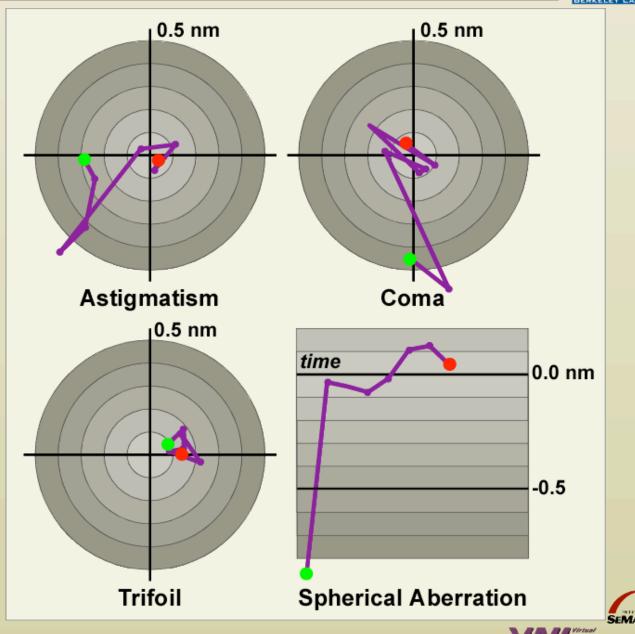
# Wavefront measurements during alignment



central field point

- Initial value
- Value following final alignment

Following initial alignment and measurement, the optic was removed and replaced in the chamber as components for imaging were installed.



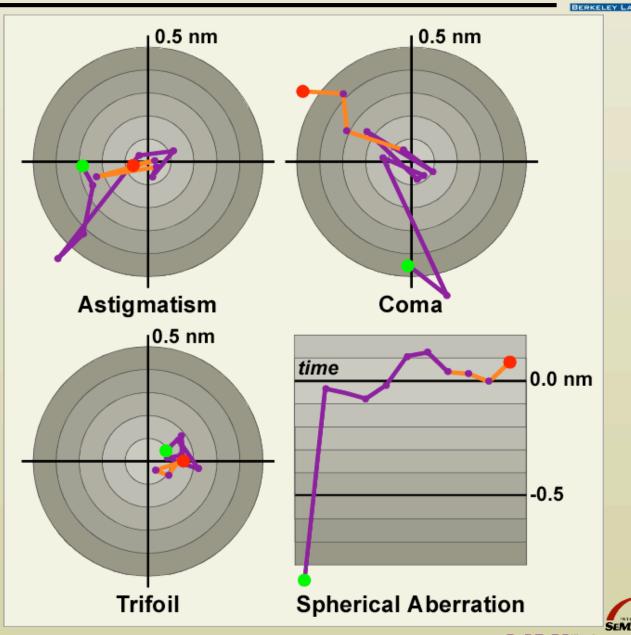
# Wavefront measurements during alignment

DERKELEY LAB

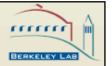
central field point

- Initial value
- Observed drift over 1 month
- Last measured value

The cause of the drift was never established



#### How sensitive is the wavefront to actuation?



Six arms support the M1 mirror.

A <u>1-µm</u> change in the arm length yields the following wavefront changes:

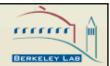
[nm]

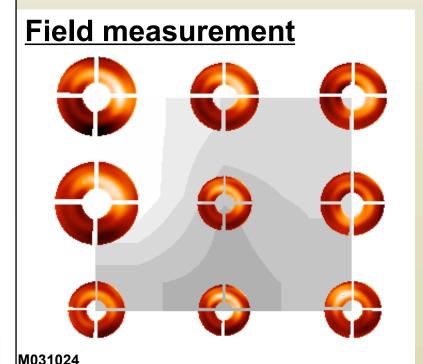
aberration coefficient	Arm 1	Arm 2	Arm 3	Arm 4	Arm 5	Arm 6
coma	3.191 57.3°	2.765 -167.8°	3.020 -76.2°	3.047 69.6°	3.323 165.3°	2.647 -44.3°
astigmatism	0.192 -119.4°	0.213 156.2°	0.161 12.2°	0.177 85.2°	0.135 -21.2°	0.044* -121.7°
spherical aberration	0.071	0.069	0.071	0.100	0.065	0.082
Δ(Wavefront)						

Mirror actuation also affects the field position



#### Final alignment state of the optic





σ <sub>37</sub> [nm]				
1.16	1.00	0.99		
1.22	0.80	0.94		
0.83	0.76	0.83		

 $\lambda / 17.8$ 

# middle-bottom field point [nm]

astig = 0.03

coma = 0.51

sph. ab. = 0.04

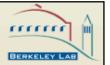
**trifoil = 0.08** 

h.-o. sph. = 0.37

One month after the final alignment, the system had drifted slightly out of its optimized alignment

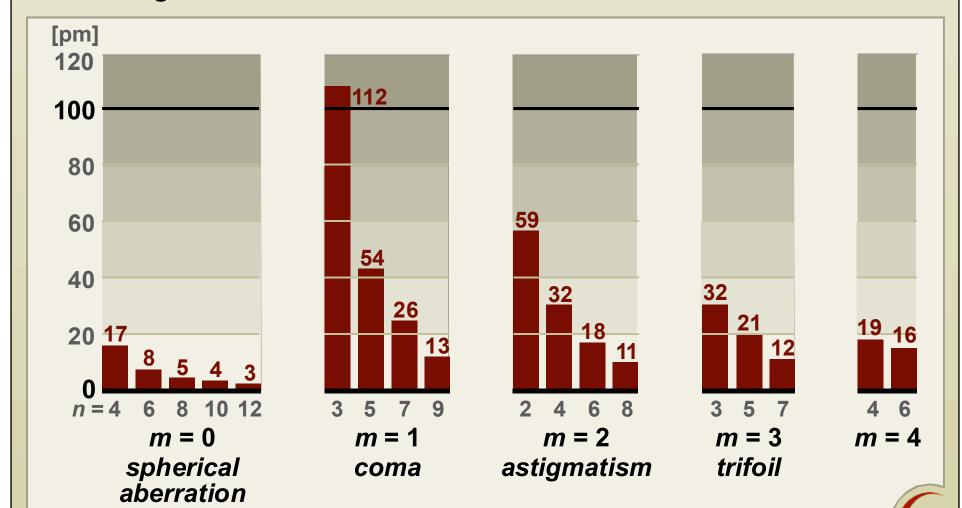


#### How precise or repeatable is shearing interferometry?

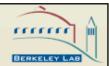


#### 1) Instantaneous repeatability

The variation of the Zernike coefficients within a set of measurements — averaged over hundreds of measurement sets.



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#### 1) Instantaneous repeatability

The variation of the Zernike coefficients within a set of measurements — averaged over hundreds of measurement sets.

#### 2) Across-the-field measurements

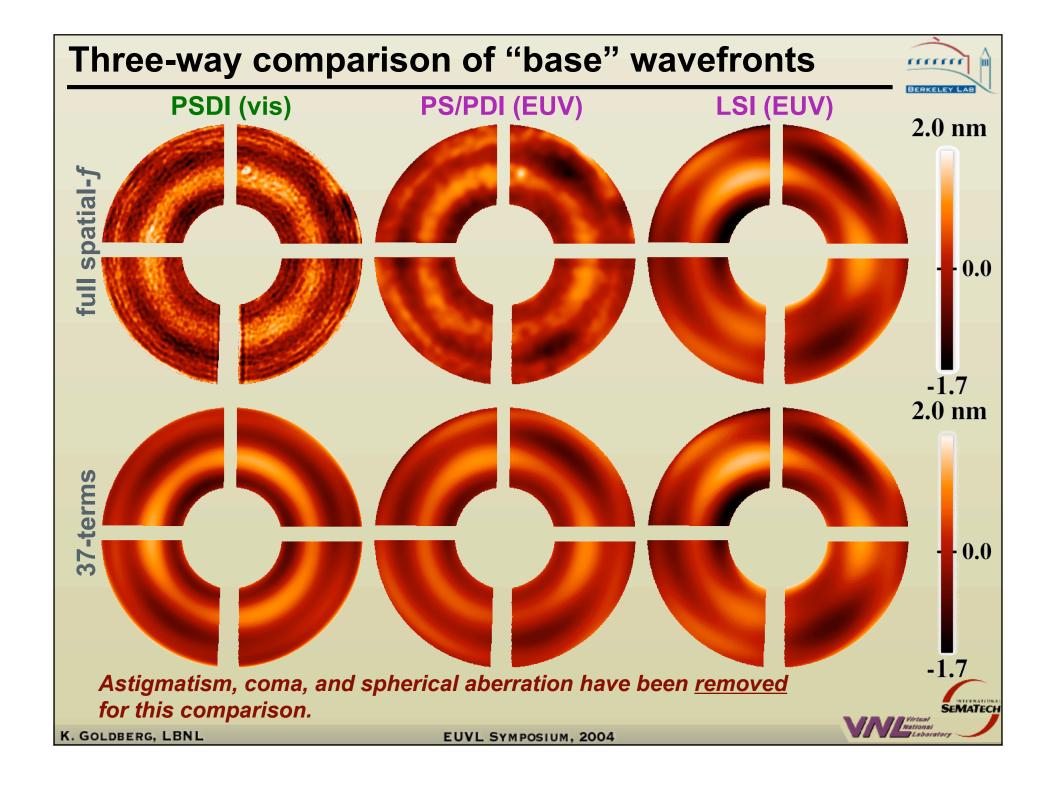
We observed small, self-consistent variations from point to point despite:

- (a) different pinholes
- (b) Over 3 mm of system travel.

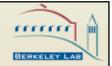
#### 3) Measurement during alignment

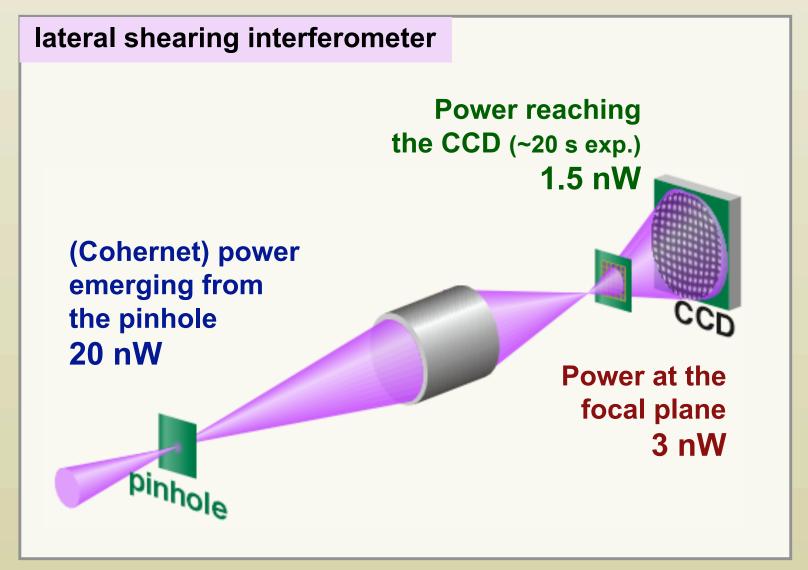
System alignment to remove astigmatism, coma, spherical aberration requires stable, self-consistent measurements. We routinely achieved ~0.05-nm control.





#### How much EUV power do you need for interferometry?







#### **Compact, Coherent EUV Source Development** at the New EUV Science & Technology Center

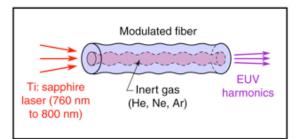


Colorado State University **Fort Collins** (Rocca, Menoni et al.)

Ni-like Ag EUV Laser 10<sup>9</sup>  $\lambda = 13.9 \text{ nm}$  $(4d \rightarrow 4p)$ Intensity (arb. units) (200 nJ @ 5 Hz)  $10^{7}$ 13.9 nm Length (mm)

**University of Colorado** Boulder (Murnane, Kapteyn et al.)

> **EUV High Harmonic** Generation (HHG)

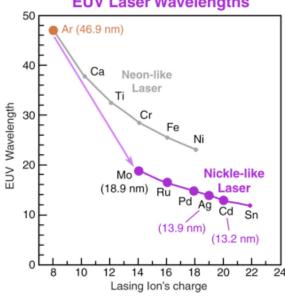


**University of California** Berkeley & LBNL (Attwood, Anderson et al.)

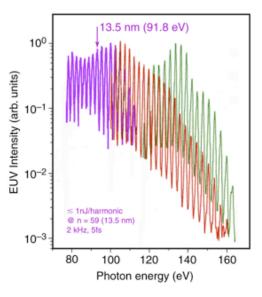
> Applications to EUV Metrologies:

- · Compact, at-wavelength **EUV** interferometry
- Compact, EUV source for defect inspection
- Compact, EUV sources for EUV microscopy
- · Compact, EUV sources for resist development

#### **EUV Laser Wavelengths**



#### **Tunable EUV Harmonics**

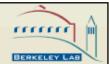


Courtesy of David Attwood



K. GOLDBERG, LBNL

#### **Conclusions**



Successful EUV interferometry at 0.3 NA.

Repeated measurements made across the field during alignment optimization.

Interferometry, alignment brought the system to diffraction-limited wavefront quality:  $\sigma_{37}$  = 0.55 nm,  $\lambda_{EUV}/24.5$ 

Alignment drift complicated measurements and comparisons.

Final wavefront at central field point:  $\sigma_{37}$  = 0.8 nm,  $\lambda_{EUV}/17$ .

Comparisons with PSDI (vis) showed consistent higher-order spherical aberration, but weak agreement in non-rotationally symmetric terms. LSI-to-PS/PDI comparison revealed subtle aspects of the data analysis that are undergoing further study.

Acknowledgment: Kim Dean, International SEMATECH

